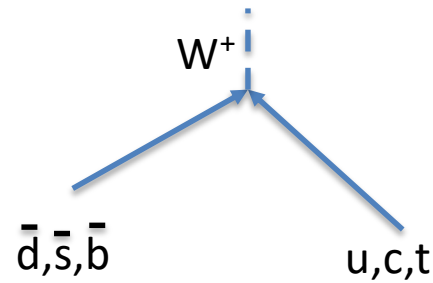
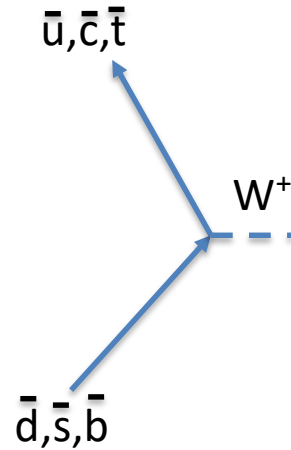
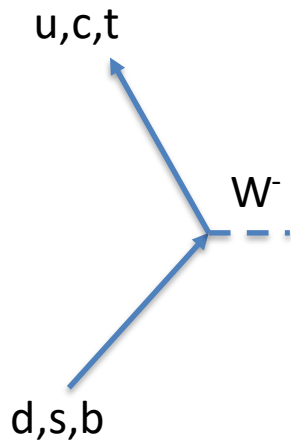


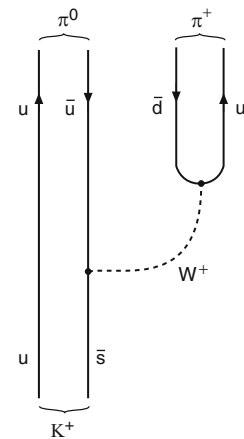
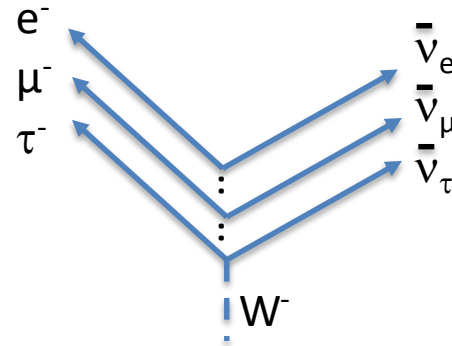
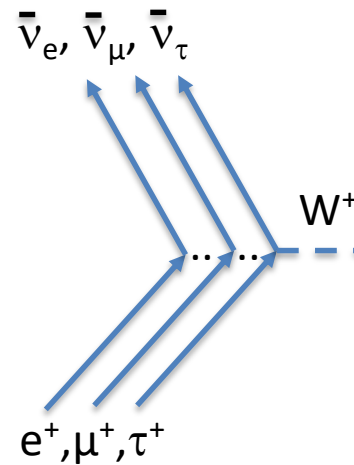
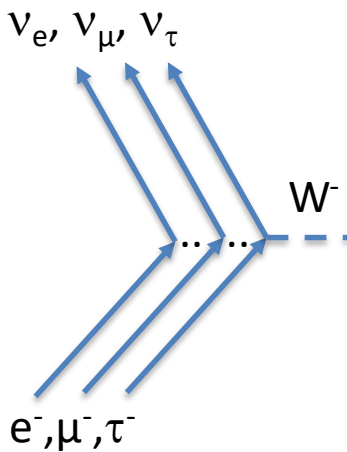
Weak Interactions



or reverse with opposite charge W

Example Reactions:

- $n \rightarrow p + e^- + \bar{\nu}_e$
- $\Lambda \rightarrow \pi^- + p$
- anti-n decay
- $pp \rightarrow D$, beta+ decay
- $\bar{\nu}_e + p \rightarrow n + e^+$
- $\pi^- \rightarrow \mu^- + \bar{\nu}_\mu$
- $\mu^- \rightarrow e^- + \nu_\mu + \bar{\nu}_e$
- $\tau^- \rightarrow \pi^- + \nu_\tau$



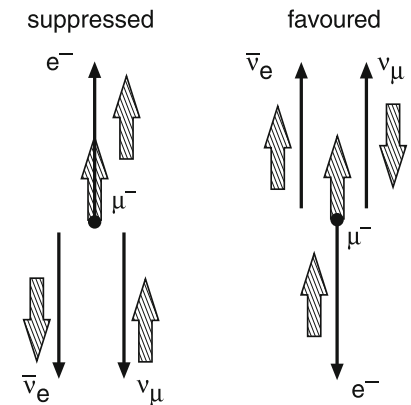
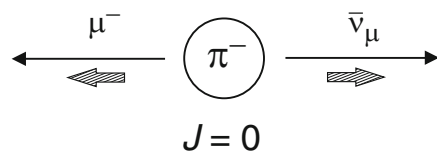
More about the Weak Interaction:

- Reminder: $\mathcal{M}_{fi} \propto g \cdot \frac{1}{Q^2 c^2 + M_W^2 c^4} \cdot g$ $\frac{G_F}{\sqrt{2}} = \frac{\pi \alpha}{2} \cdot \frac{g^2}{e^2} \cdot \frac{(\hbar c)^3}{M_W^2 c^4}$
- Violates Parity (V-A structure)
 - For hadrons: Have separate couplings $G_V = G_F$, G_A
- Quark mixing: Mass eigenstates $\neq$ participants in the weak interaction: $|d'\rangle = \cos \theta_C |d\rangle + \sin \theta_C |s\rangle$
 (Cabibbo angle) $|s'\rangle = \cos \theta_C |s\rangle - \sin \theta_C |d\rangle$,
 $\sin \theta_C \approx 0.22$, and $\cos \theta_C \approx 0.98$
- Cabibbo-Kobayashi-Maskawa matrix:

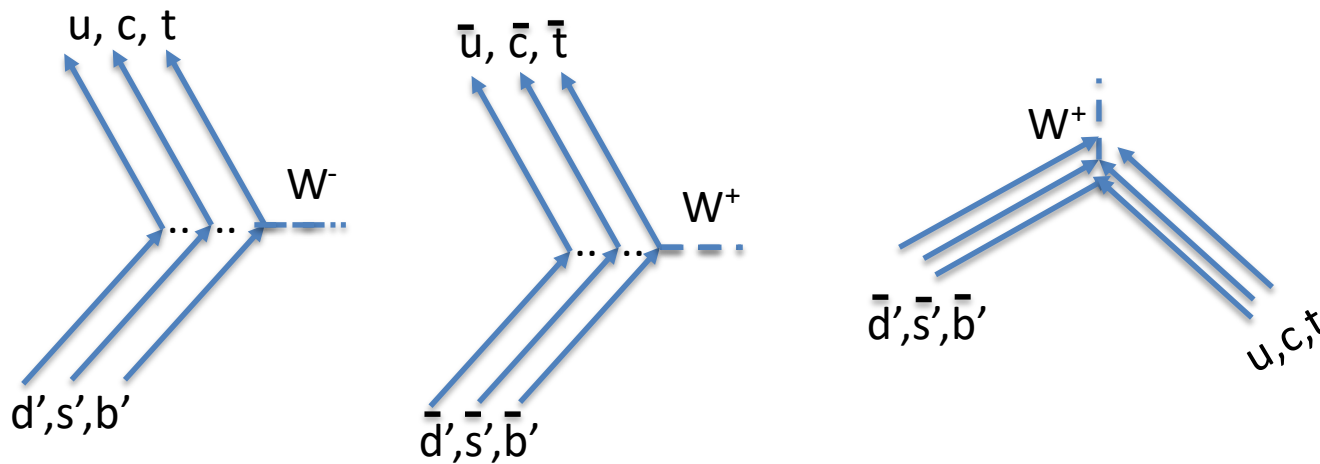
$$\begin{pmatrix} |d'\rangle \\ |s'\rangle \\ |b'\rangle \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \cdot \begin{pmatrix} |d\rangle \\ |s\rangle \\ |b\rangle \end{pmatrix} \quad (|V_{ij}|) = \begin{pmatrix} 0.974 & 0.225 & 0.003 \\ 0.225 & 0.973 & 0.041 \\ 0.008 & 0.040 & 0.999 \end{pmatrix}$$

Parity Violation

- Maximal in the weak interaction!
- W 's couple only to “left-handed” particles
- Explains preference for μ in π decay



Weak Interactions with Weak Eigenstates



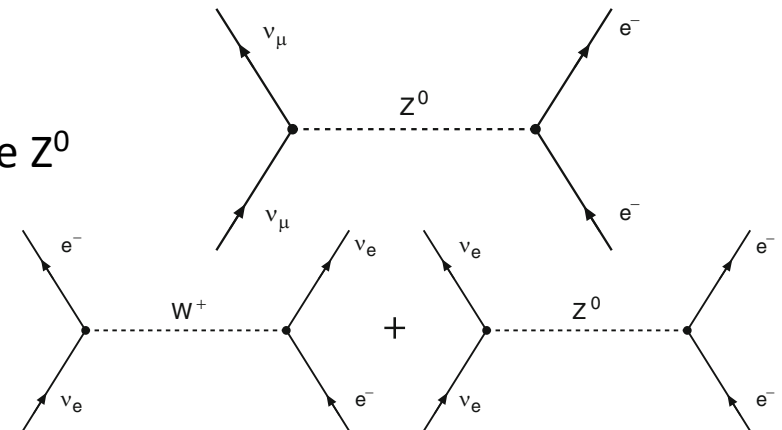
or reverse with opposite charge W

Why don't the Lepton States mix?

THEY DO! But we cannot observe the mass eigenstates of the neutrinos. See later...

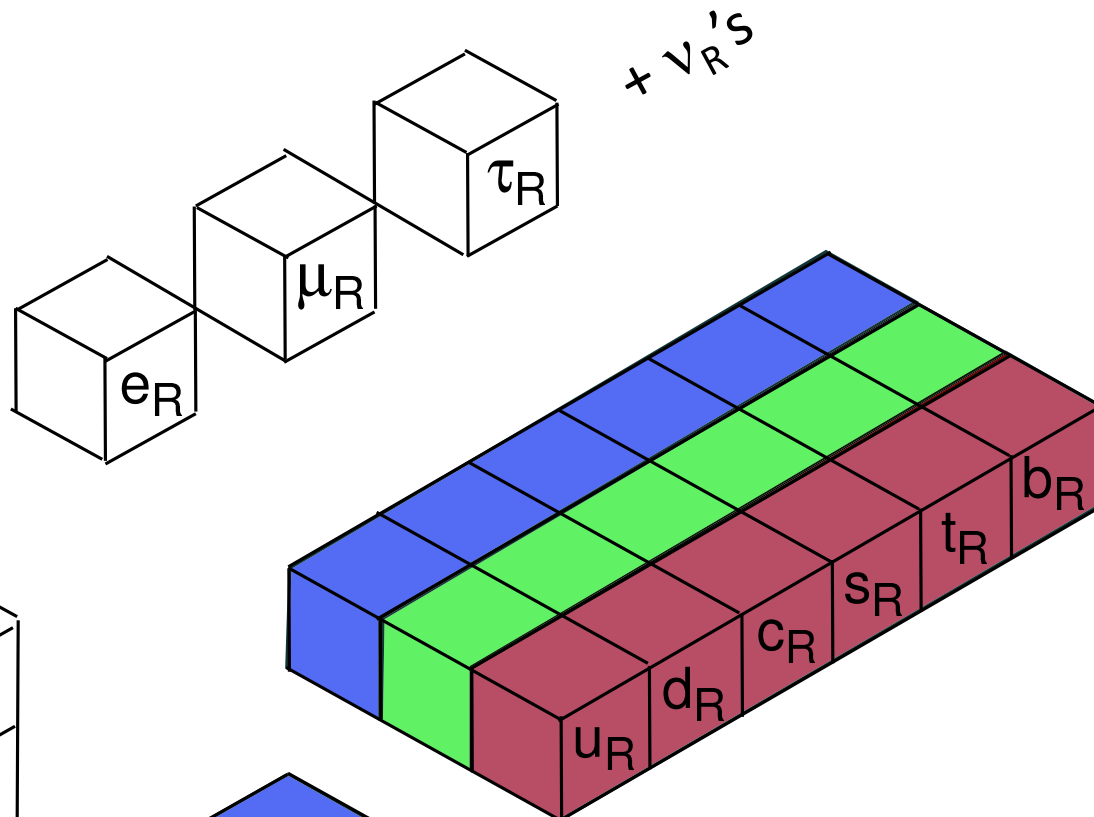
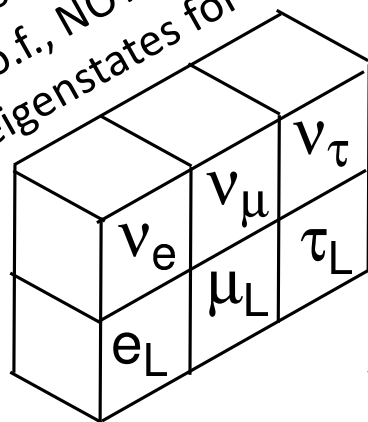
Why is there no W^0 ?

THERE IS! However, in "real life" we instead observe the Z^0 which is slightly more massive than the W 's. see later...



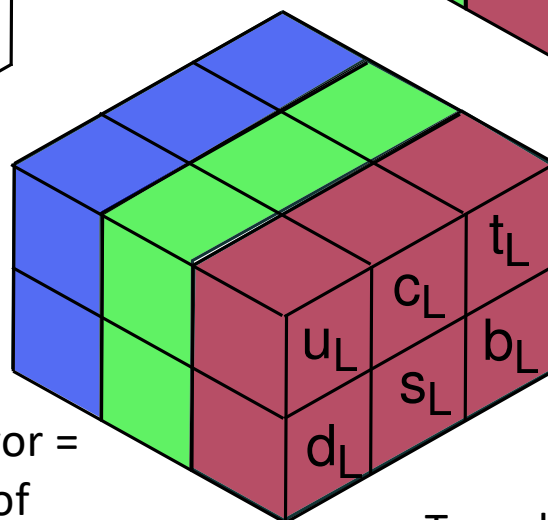
$+ \nu_R$'s

These are the EW
d.o.f., NOT the mass
eigenstates for ν 's

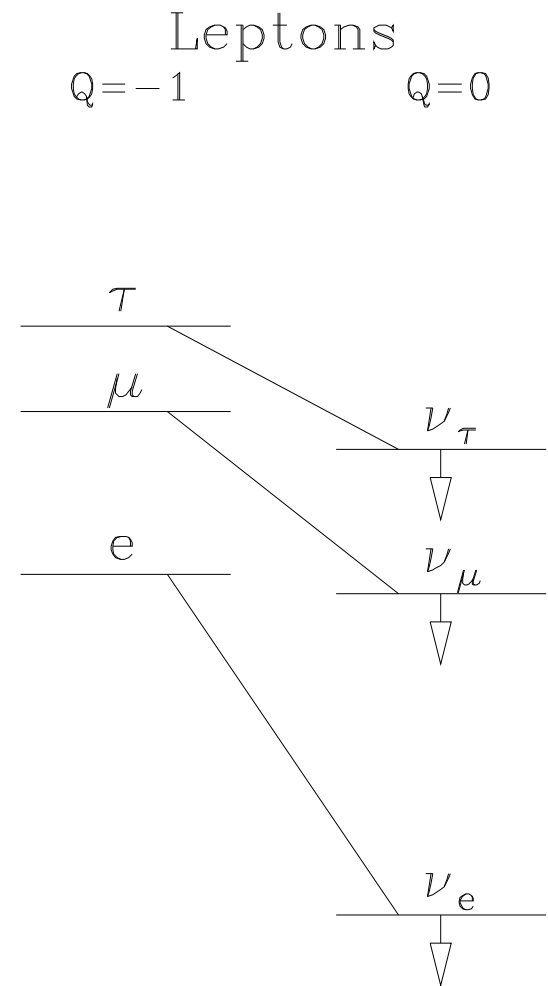
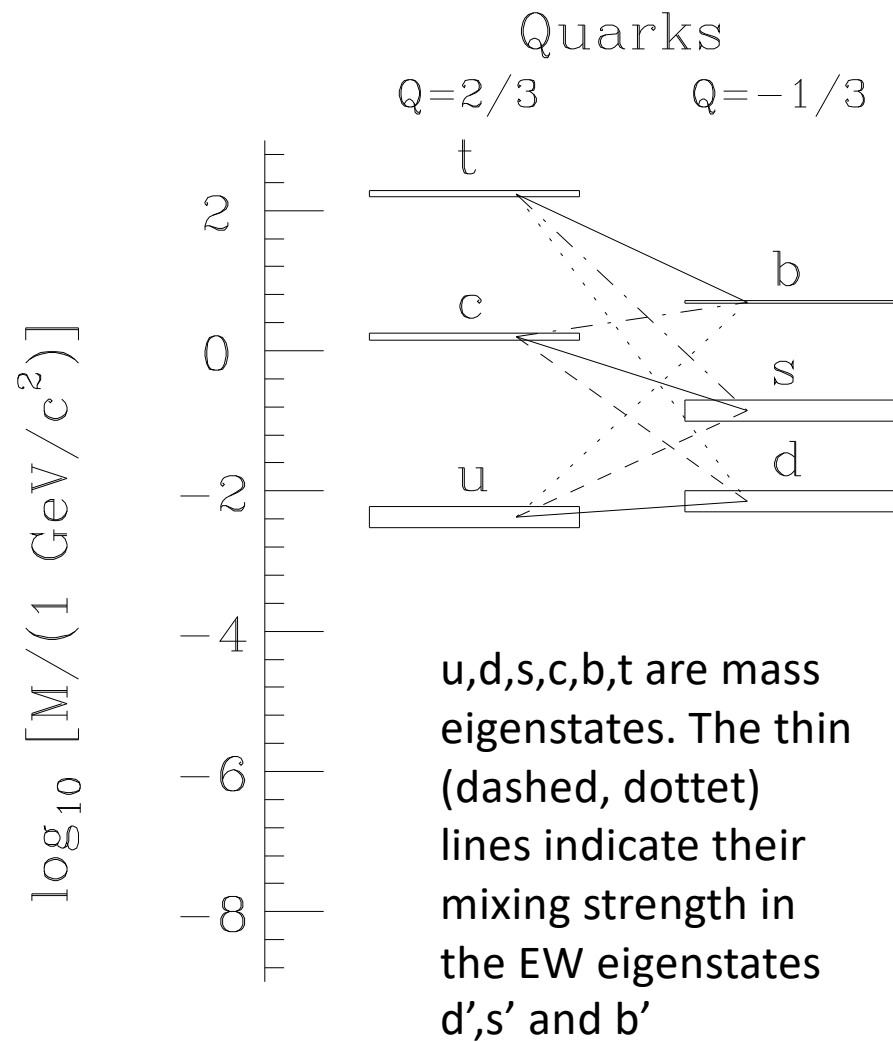


Repeat with
antiparticles but
exchange R vs. L

These are the Flavor =
mass eigenstates of
the quarks



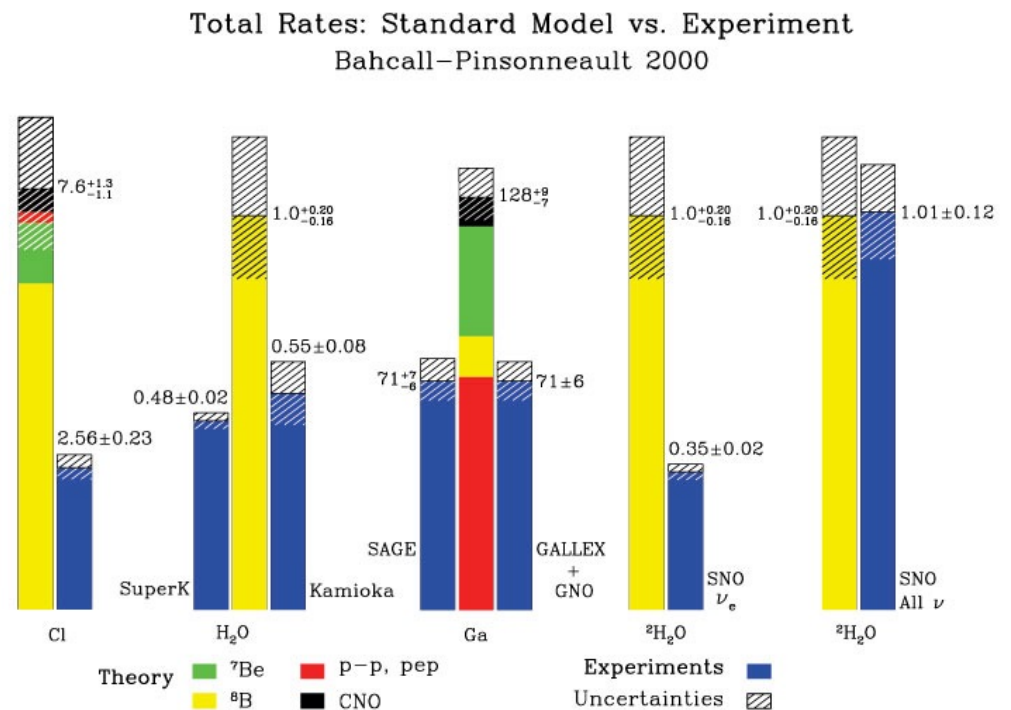
True d.o.f. for EW theory: d', s', b'



The neutrinos are the EW eigenstates. They are NOT mass eigenstates.

Neutrinos DISAPPEAR!

- Originally discovered by Ray Davis: there are too few neutrinos coming from the sun
- Original experiment in Homestead Mine (Cl): Only 1/3 of expected flux
- Confirmed by Sage, Gallex, Super-K, SNO, ...
- Confirmed with reactors: Bugey, Chooz, KamLand,... and accelerator neutrinos (T2K, NOvA,...)
- Also found disappearance of μ -neutrinos in atmosphere: Super-K.



Explanation: 2 –neutrino model

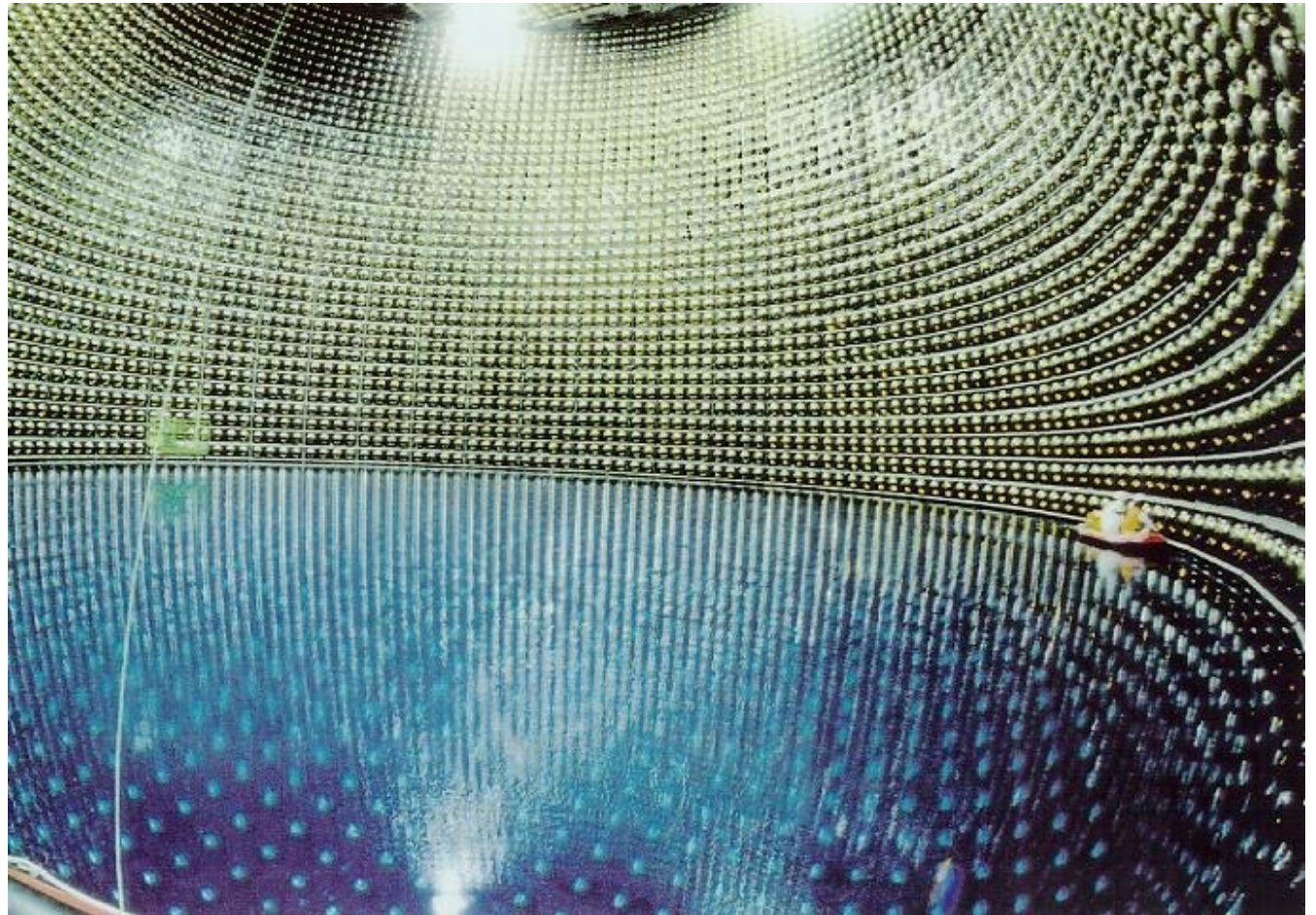
$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2\left(\frac{\Delta m^2 L}{4E}\right) \text{ (natural units).}$$

$$P_{\alpha \rightarrow \beta, \alpha \neq \beta} = \sin^2(2\theta) \sin^2\left(1.27 \frac{\Delta m^2 L [\text{eV}^2] [\text{km}]}{E [\text{GeV}]}\right).$$

- The mass differences, Δm^2 , are known to be on the order of $1 \times 10^{-4} \text{ eV}^2$
- Oscillation distances, L , in modern experiments are on the order of [kilometers](#)
- Neutrino energies, E , in modern experiments are typically on order of MeV or GeV.

Kamiokande, Super-K

- Detect neutrinos from sun and atmospheric neutrinos
- Only 50% of solar vs
- Detection via Cherenkov Light



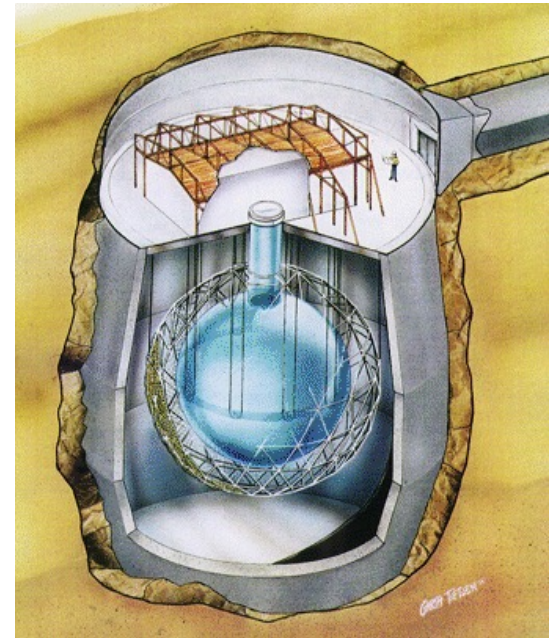
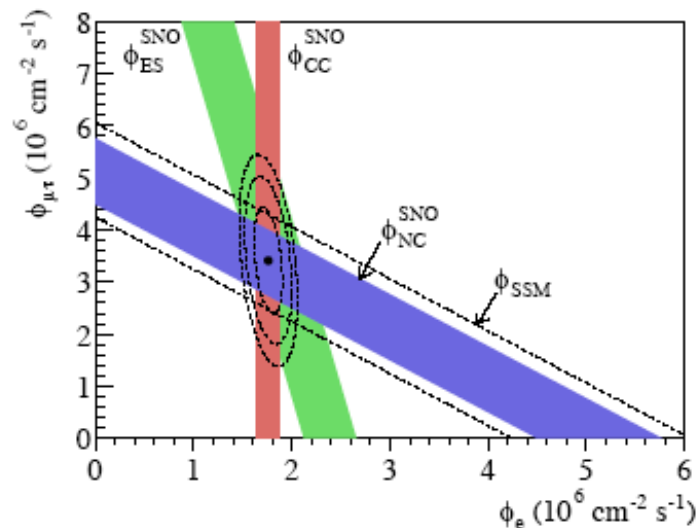
Kamioka Observatory, ICRR (Institute for Cosmic Ray Research), The University of Tokyo

SNO

- Heavy Water Cherenkov detector
- Sensitive to all 3 types of ν 's with different observables:

$$d + \nu_e \rightarrow p + p + e^-;$$

$$d + \nu_\mu \rightarrow p + n + \nu_\mu$$
- First unambiguous confirmation that total number of ν 's from sun is as expected - only flavor changes





The Nobel Prize in Physics 2015
Takaaki Kajita, Arthur B. McDonald

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The Nobel Prize in Physics 2015



Photo © Takaaki Kajita

Takaaki Kajita

Prize share: 1/2

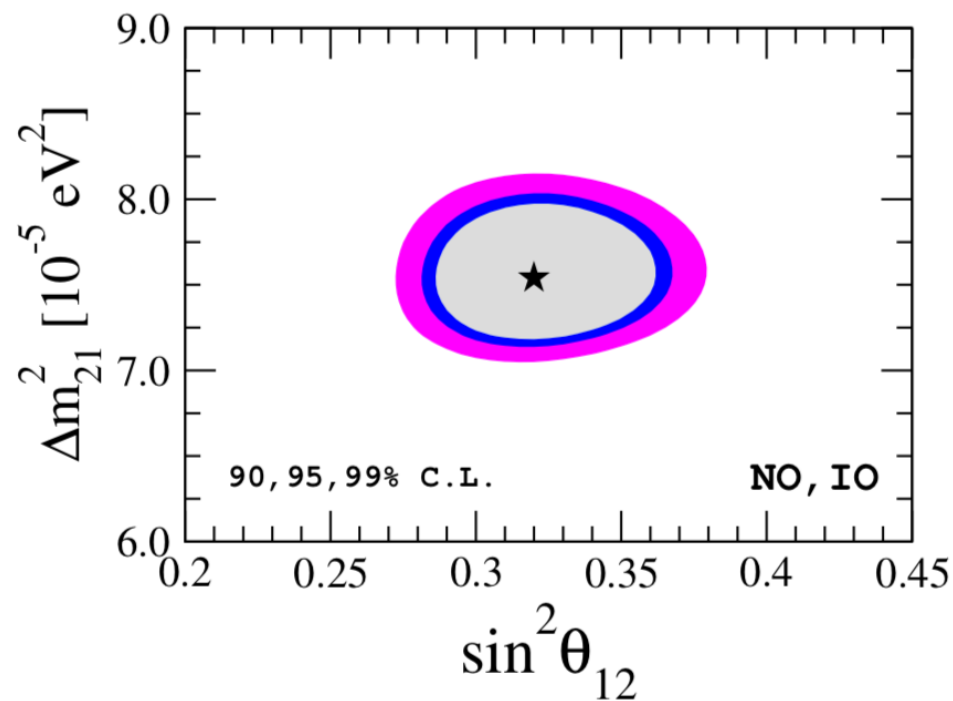
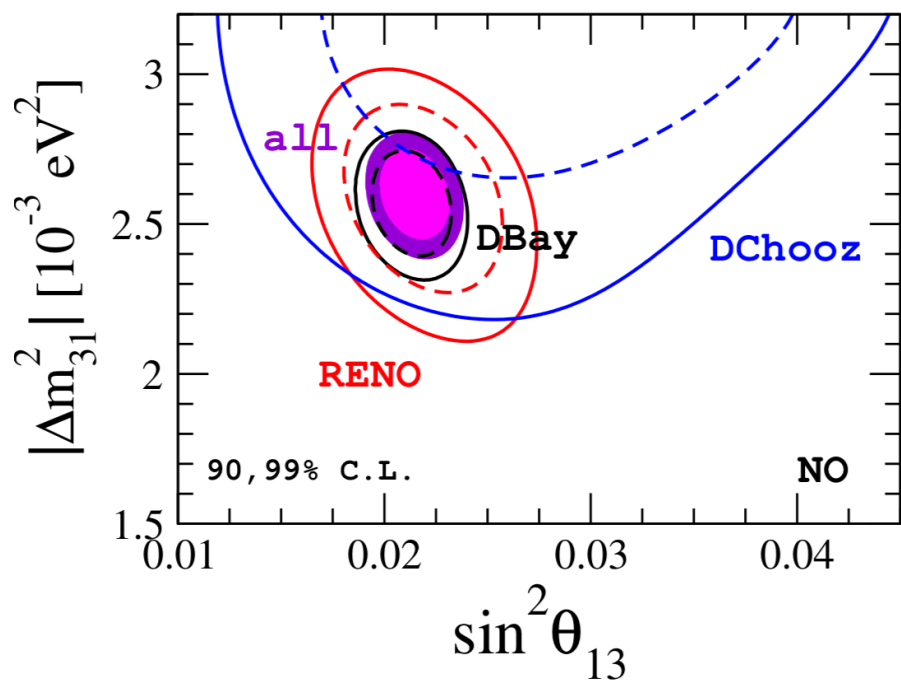
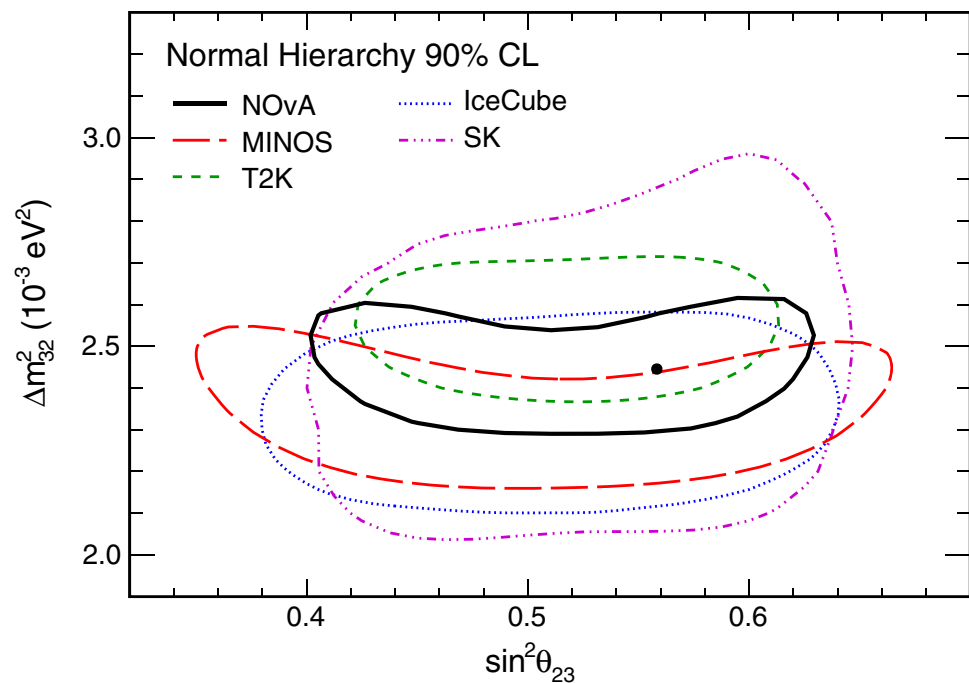
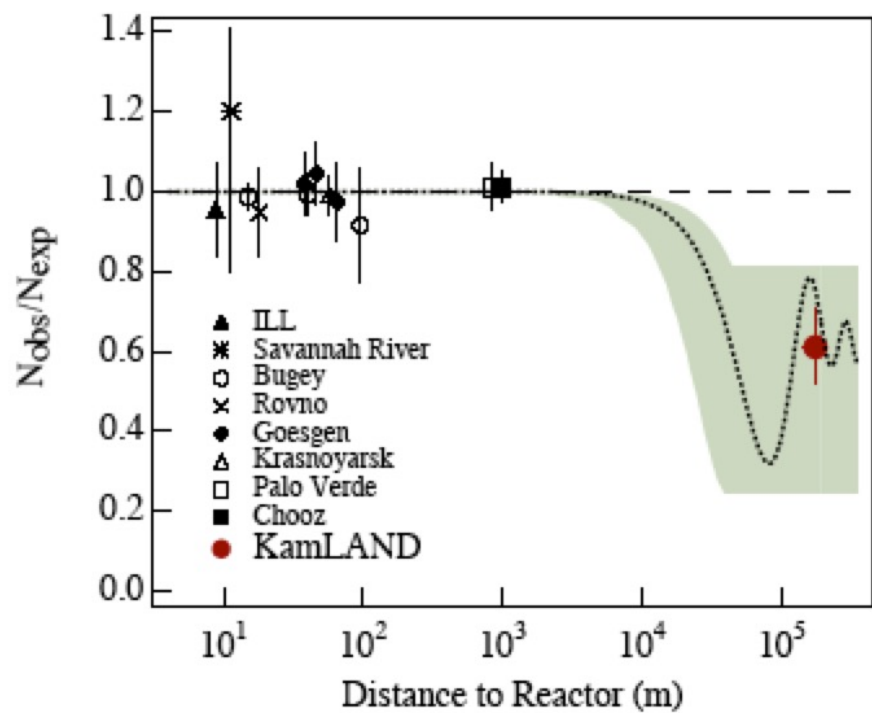


Photo: K. MacFarlane.
Queen's University
/SNOLAB

Arthur B. McDonald

Prize share: 1/2

The Nobel Prize in Physics 2015 was awarded jointly to Takaaki Kajita and Arthur B. McDonald *"for the discovery of neutrino oscillations, which shows that neutrinos have mass"*



Electroweak Interaction

(Unification by Salam and Weinberg *et al.*)

- Just like the strong and EM interaction, Weak interactions are mediated by “charges”
 - Named “Hypercharge Y ” and “Weak Isospin I ”
 - W^+ , W^- , W^0 and couple with strength g to I
 - B^0 couples with strength g' to Y
 - $g \sin \theta_W = g' \cos \theta_W = e. \quad \sin^2 \theta_W = 0.23113 \pm 0.00021$
- Huh? Because of their interaction with the Higgs field, some particles we actually observe

are mixtures: $|\gamma\rangle = \cos \theta_W |B^0\rangle + \sin \theta_W |W^0\rangle$

$$\frac{M_W}{M_Z} = \cos \theta_W \approx 0.88 \quad |Z^0\rangle = -\sin \theta_W |B^0\rangle + \cos \theta_W |W^0\rangle$$

Higgs Field

- Create “Drag” on Particles (“Molasses”)
- *) Origin of Mass
Makes some gauge bosons very heavy (W 's, Z 's) and therefore short-range (“Weak” interaction)
- Origin of electroweak symmetry breaking
- Pointlike, Fundamental

POPULAR ANALOGIES FOR THE HIGGS FIELD



IT'S LIKE MOVING
THROUGH TREACLE
(OR MOLASSES)



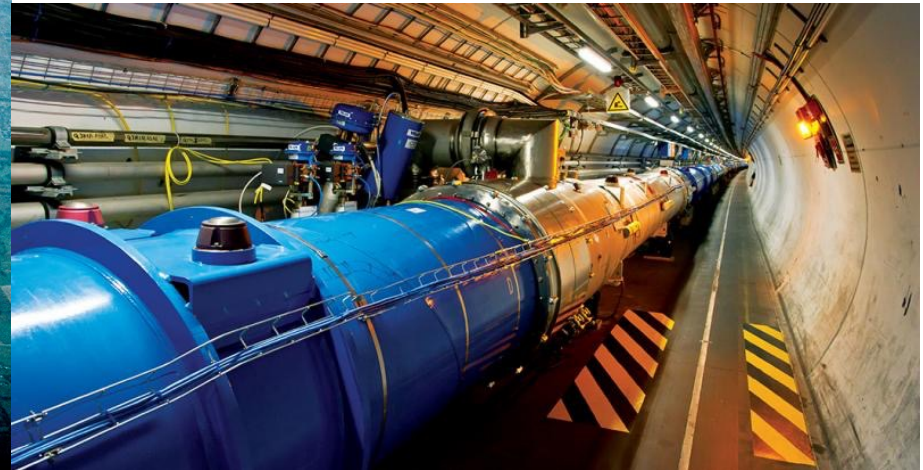
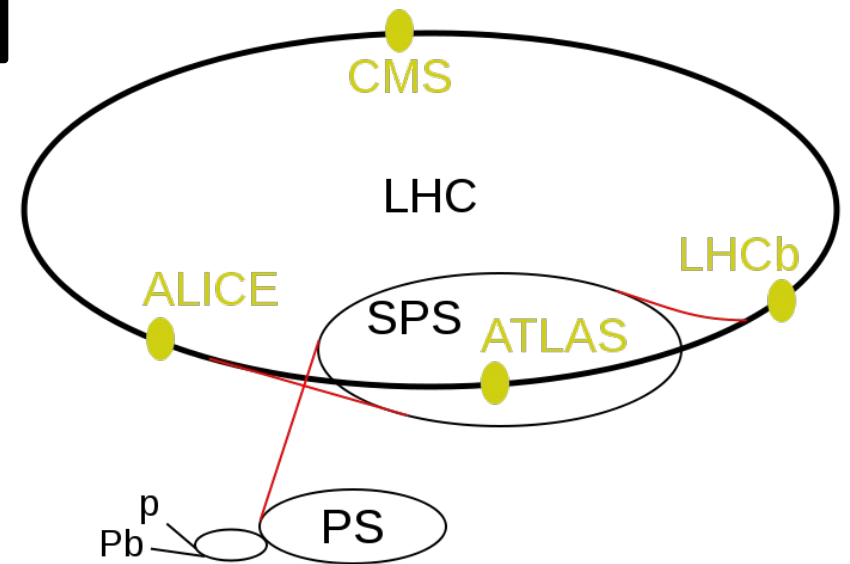
IT'S LIKE A POLITICIAN
MOVING THROUGH A
CROWD



IT'S LIKE MOVING
THROUGH A CROWD
OF POLITICIANS
COVERED IN TREACLE



The LHC at CERN



- See also the movie “Particle Fever”

Higgs Discovery

