

- 1) Use the web page [www.nndc.bnl.gov/wallet/](http://www.nndc.bnl.gov/wallet/) ( or equivalent) to find the energy release in the following fusion reactions.
  - a.  ${}^4\text{He} + {}^{12}\text{C} \rightarrow {}^{16}\text{O}$
  - b.  ${}^{12}\text{C} + {}^{12}\text{C} \rightarrow {}^{24}\text{Mg}$
  - c.  $p + {}^{12}\text{C} \rightarrow {}^{13}\text{N} + \gamma$ .

- 2) **Astrophysics in a Nutshell** Chapter 3 Problem 7.

The nuclear reaction rate in a star is proportional to

$$\langle \sigma v \rangle \propto \int_0^{\infty} f(E) dE$$

$$f(E) = e^{-E/(kT)} e^{-\sqrt{E_G/E}}$$

with  $E_G = (\pi\alpha)^2 M_p c^2$ ,  $\alpha = e^2 / (\hbar c) = 1/137...$

- a. Show that the maximum value of  $f(E)$  occurs at

$$E = E_0 = [kT/2]^{2/3} E_G^{1/3}$$

- b. Form a Taylor series, to second order in  $\ln[f(E)]$ , to approximate  $f(E)$  as a Gaussian; i.e. find  $A$  and  $\Delta$  in the approximation

$$f(E) \approx A e^{-(E-E_0)^2 / (2\Delta^2)}$$

- c. Using the Gaussian identities

$$\int_{-\infty}^{\infty} e^{-x^2/2} dx = \sqrt{2\pi} = \int_{-\infty}^{\infty} x^2 e^{-x^2/2} dx, \text{ show that}$$

$$\int_0^{\infty} f(E) dE \approx \sqrt{2\pi} f(E_0) \Delta \text{ if } E_0 \gg \Delta.$$

- 3) The pp chain produces 2 neutrinos for every 26.2 MeV of fusion energy release. The solar flux of visible light reaching the upper atmosphere of the earth is  $\approx 1000 \text{W/m}^2$ . Assume the typical photon in the solar spectrum is a green photon of energy  $h\nu = 2 \text{eV}$ .

- a. What is the number flux of visible photons reaching the upper atmosphere of the earth?
- b. What is the number flux of neutrinos reaching the earth?
- c. If the average energy of these neutrinos is 200 KeV, what fraction of the solar luminosity is carried away by neutrinos?

4) **Astrophysics in a Nutshell** Chapter 3 Problem 8.

The power production per unit mass of the  $pp \rightarrow \text{De}^+ \nu$  reaction is (3.134)

$$\epsilon = \frac{\rho}{M_H} \frac{2^{2/3}}{\sqrt{3}} \frac{QS_0c}{M_H \sqrt{M_H c^2}} \frac{E_G^{1/6}}{(kT)^{2/3}} e^{-3[E_G/(4kT)]^{1/3}}$$

The “astrophysical S-factor” is  $S_0 = 4 \cdot 10^{-46} \text{ cm}^2 \text{ KeV}$ . For the first step in the chain,  $Q = 2.2 \text{ MeV}$ , but for the whole pp chain  $Q = 26.2 \text{ MeV}$ . Use values from the solar center,  $\rho = 150 \text{ g/cm}^3$ .

- a. Perform dimensional analysis on the expression for  $\epsilon$  to establish it has units of Energy per unit mass per second.
- b. Approximate  $\epsilon$  as a power law  $\epsilon \approx \epsilon_0 (T/T_0)^\beta$ . Hint, make a Taylor series in  $\ln \epsilon$  as a function of  $\ln(T/T_0)$  and expand around  $\ln(T/T_0) = 0$ .
- c. Evaluate  $\beta$  for  $kT_0 = 1.0 \text{ KeV}$  and  $E_G = 500 \text{ KeV}$ .